

10/12/16 Lecture 6 outline

- Momentum generates spatial translations, and the Hamiltonian (energy) generates time translations. Schrodinger equation:  $i\hbar\partial_t U(t, t_0) = HU(t, t_0)$ . For time independent  $H$ , this is simple to integrate:  $U = e^{-iH(t-t_0)/\hbar}$ . For  $t$  dependent  $H$ , one gets instead (Dyson)  $U(t, t_0) = T \exp(-i/\hbar \int_{t_0}^t dt' H(t'))$ , where  $T$  denotes time ordering. We won't need that here. Since time evolution is a unitary operator, it preserves the norm of kets.

- Schrodinger picture has the time dependence in the bras and kets:  $|\psi(t)\rangle_S = U(t, t_0)|\psi(t_0)\rangle$ . So  $|\psi(t)\rangle_S$  satisfies the Schrodinger equation:  $i\hbar\partial_t|\psi(t)\rangle_S = H|\psi(t)\rangle_S$ . The Heisenberg picture is equivalent, but instead puts the  $t$  dependence in the operators:

$$\langle \mathcal{O} \rangle = {}_S\langle \psi(t) | \mathcal{O}_S | \psi(t) \rangle_S = \langle \psi(t_0) | U(t, t_0)^\dagger \mathcal{O} U(t, t_0) | \psi(t_0) \rangle \equiv \langle \psi(t_0) | \mathcal{O}^H(t) | \psi(t_0) \rangle,$$

where  $\mathcal{O}^H = U^\dagger \mathcal{O} U$  is the Heisenberg picture operator. Note that it satisfies

$$\frac{d}{dt} \mathcal{O}^H = \frac{1}{i\hbar} [\mathcal{O}^H, H] + \frac{\partial}{\partial t} \mathcal{O}^H.$$

This is exactly what one finds in classical mechanics, in terms of Poisson brackets, with Dirac's replacement  $[A, B]_{QM} = i\hbar[A, B]_{PB}$ . For example, if  $H = \vec{p}^2/2m + V(\vec{x})$ , the above gives

$$\frac{d}{dt} \vec{x}^H(t) = \vec{p}^H(t)/m, \quad \frac{d}{dt} \vec{p}^H(t) = -\partial_{\vec{x}} V(\vec{x}).$$

These look like the classical equations of motion in the Hamiltonian formulation; they apply as operator statements to the Heisenberg picture operators.

- Free particle: Heisenberg picture operators have  $\vec{p}(t) = \vec{p}(0) = \vec{p}$  and  $\vec{x}(t) = \vec{x}_0 + \vec{p}t/m$ ; these are all operators. So  $[x_i(t), x_i(0)] = -i\hbar t/m$ , and thus  $\langle (\Delta x_i(t))^2 \rangle \langle (\Delta x_i(0))^2 \rangle \geq \hbar^2 t^2 / 4m^2$ .

- E.g. if there is an external magnetic field in between the SG experiments, it'll make the spin of the state precess via  $H = -\vec{\mu} \cdot \vec{B}$  with  $\vec{\mu} = ge\vec{S}/2mc$ . Take e.g.  $\vec{B} = B_0 \hat{z}$  so  $H = g|e|S_z B / 2mc \equiv \omega S_z$ . Then  $U = e^{-iHt/\hbar}$  is diagonal in the  $|\pm_z\rangle$  basis. Show e.g. that if in the  $|+_x\rangle$  state at  $t_0 = 0$ , the probability of finding it later in the  $|\pm_x\rangle$  state is  $\cos^2(\omega t/2)$  and  $\sin^2(\omega t/2)$ , respectively, and  $\langle S_x \rangle = \frac{1}{2}\hbar \cos \omega t$  and  $\langle S_y \rangle = \frac{1}{2}\hbar \sin \omega t$  and  $\langle S_z \rangle = 0$ , fitting with the classical picture of precessing in the  $xy$  plane with frequency  $\omega$ . Recall from HW that  $e^{i\theta \hat{n} \cdot \vec{\sigma}} = \cos \theta \mathbf{1} + i \sin \theta \hat{n} \cdot \vec{\sigma}$ . Work out  $S_x(t)$  in that basis.