4/10/17 Lecture 3 outline

• Last time:  $S = -mc^2 \int d\tau$  and L for relativistic particle. Relativistic charged particle  $S \supset -(q/c) \int A_{\mu} dx^{\mu}$  and term in  $L \supset -q\phi + (q/c)\vec{v} \cdot \vec{A}$ . Canonical momentum  $\vec{p} = \partial L/\partial \vec{v} = \gamma m \vec{v} - q \vec{A}/c$  and  $H = \vec{p} \cdot \vec{v} - L = \gamma mc^2 - q\phi$  (magnetic fields do no work). But H should be expressed in terms of  $\vec{p}$ .

• QM: replace  $\hat{p}^{\mu} \to i\hbar\partial^{\mu}$  in position space. Consider the S.E.

• Punchline: in non-rel limit get  $i\hbar D^0\psi = (\hbar^2/2m)\vec{D}^2\psi$ , where  $D^{\mu} \equiv (D^0, \vec{D}) = \partial^{\mu} - (q/i\hbar c)A^{\mu}$ .

Get  $p^{\mu} \to p^{\mu} - (q/c)A^{\mu}$  when a charged particle is in an  $\vec{E}$  and  $\vec{B}$  field. For QM,  $p^{\mu} \to i\hbar\partial^{\mu}$  in position space, so get e.g.  $i\hbar D^{0}\psi = (-\hbar^{2}/2m)\vec{D}^{2}\psi$ , where (punchline)  $D^{\mu} \equiv (D^{0}, \vec{D}) = \partial^{\mu} - (q/i\hbar c)A^{\mu}$  is the covariant derivative.

•  $S = \ldots - q/c \int A_{\mu} dx^{\mu}$  and gauge invariance.

• Gauge transformations  $A_{\mu} \to A_{\mu} + \partial_{\mu} f(x)$ , and local U(1) phase rotation of  $\psi \to e^{-iq/\hbar c}\psi$ , with  $D_{\mu}\psi \to e^{-iq/\hbar c}D_{\mu}\psi$ .

• Hamiltonian picture:  $i\hbar\partial_t U(t,t_0) = HU(t,t_0), i\hbar|\psi(t)\rangle_S = H|\psi(t)\rangle_S$ , and  $\mathcal{O}^H = U^{\dagger}\mathcal{O}U$  has

$$\frac{d}{dt}\mathcal{O}^{H} = \frac{1}{i\hbar}[\mathcal{O}^{H}, H] + \frac{\partial}{\partial t}\mathcal{O}^{H}.$$

Consider e.g. the SHO,  $H = p^2/2m + \frac{1}{2}m^2\omega^2 x^2$ . The equation  $H|n\rangle = E_n|n\rangle$  in position space becomes a 2nd-order differential equation which has solution given by some special functions. Happily, there is a much better way to solve this problem, which is simpler, more interesting, and more important than solving a differential equation. It uses creation and annihilation operators

$$a \equiv \sqrt{\frac{m\omega}{2\hbar}} (x + ip/m\omega),$$
 so  $a^{\dagger} \equiv \sqrt{\frac{m\omega}{2\hbar}} (x - ip/m\omega).$ 

These satisfy the fundamental property  $[a, a^{\dagger}] = 1$ . We can then immediately show that the Hermitian operator  $N = a^{\dagger}a$  has eigenvalues n = 0, 1, 2..., and the eigenvectors  $|n\rangle$ satisfy  $a|n\rangle = \sqrt{n}|n-1\rangle$  and  $a^{\dagger}|n\rangle = \sqrt{n+1}|n\rangle$ . Since  $H_{SHO} = \hbar(\omega N + \frac{1}{2})$ , we're done. If we really want  $\langle x|n\rangle$ , we can get it from  $|n\rangle = (n!)^{-1/2}(a^{\dagger})^n|0\rangle$  by replacing  $p \to -i\hbar \frac{d}{dx}$ and we can solve for  $\psi_0(x) = \langle x|0\rangle$  by using  $a|0\rangle = 0$ , which in position space becomes a simple first-order differential equation for  $\psi_0(x)$ :

$$\psi_n(x) = \frac{(m\omega/2\hbar)^{n/2}}{\sqrt{n!}} (x - \frac{\hbar}{m\omega} \frac{d}{dx})^n \psi_0(x), \qquad (x + \frac{\hbar}{m\omega} \frac{d}{dx}) \psi_0(x) = 0.$$

In the Heisenberg picture we have  $\dot{a} = -i\omega a$ , hence  $a(t) = e^{-i\omega t}a$ , where a = a(0).

• Path integral description of QM, solenoids and observability of flux inside. Dirac's magnetic monopoles and quantization rule.

• Klein Gordon theory, SHO, and charged Klein Gordon theory. Covariant derivatives and minimal substitution. Euler Lagrangian equations for field theory.

•  $S = \int d^4x (-1/4) F_{\mu\nu} F^{\mu\nu} - j^{\mu} A_{\mu}/c.$ 

• Dirac equation and Lagrangian