## 5/10/18 Lecture outline

 $\star$  Reading: Zwiebach chapter 8

• Symmetries and conservation laws. Recall charge conservation  $\partial_{\mu}j^{\mu} = 0$ , which is required by gauge invariance of  $\mathcal{L} \supset A_{\mu}j^{\mu}$ , i.e.  $\delta \mathcal{L} = 0$  under  $\delta A_{\mu} = \partial_{\mu}f$ . Show that it implies conservation of  $Q = \int d^3x j^0$ .

• Recall Noether's theorem for  $L(q, \dot{q})$ : if continuous symmetry  $\delta q_i$  then  $p_i \delta q_i$  is conserved.

• Likewise, for  $S = \int d\xi^0 \dots d\xi^k \mathcal{L}(\phi^a, \partial_\alpha \phi^a)$ , a symmetry  $\delta \phi^a$  implies a conserved current  $j^\alpha = \frac{\partial \mathcal{L}}{\partial (\partial_\alpha \phi^a)} \delta \phi^a$ : show it satisfies  $\partial_\alpha j^\alpha = 0$ , so  $Q = \int d\xi^1 \dots d\xi^k$  has  $\frac{d}{d\xi^0} Q = 0$ .

For a string  $S = \int d\xi^0 d\xi^1 \mathcal{L}(\partial_\alpha X^\mu)$  (has translation invariance,  $\delta X^\mu = \epsilon^\mu$  so there is a conserved current  $\epsilon^\mu j^\alpha_\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\alpha X^\mu)} \delta X^\mu$ . So get conservation of  $j^a_\mu = \mathcal{P}^a_\mu$  (where  $a = \sigma, \tau$ ) is the conserved Noether current for spacetime translation invariance,  $\delta X^\mu = \epsilon^\mu$ . The string equations of motion are equivalent to the worldsheet conservation of this current:  $\partial_a j^a_\mu = 0$ . The spacetime momentum of the string is the corresponding conserved charge:  $p_\mu = \int d\sigma \mathcal{P}^\tau_\mu$ . So  $\frac{dp_\mu}{d\tau} = -\int_0^{\sigma_1} \partial_\sigma \mathcal{P}^\sigma_\mu = -\mathcal{P}^\sigma_\mu |_0^{\sigma_1}$ . It is conserved for the closed string, or open Neumann BCs. Not conserved for Dirichlet BCs. The Dirichlet case means that the string ends on a D-brane, and momentum can go through the string into the D-brane (their total momentum is conserved). Same for wave on a string with the ends tied down, e.g. a traveling wave is reflected, which flips  $p \to -p$ , but the difference in momentum is transferred to the post at the end and total momentum of the system is conserved.

• Note that  $p_{\mu}$  is a conserved *worldsheet* charge. It becomes a conserved spacetime charge in static gauge,  $\tau = t$ . We can write more generally the conserved flux of worldsheet current as  $(\mathcal{P}^{\tau}_{\mu}, \mathcal{P}^{\sigma}_{\mu}) \cdot (d\sigma, -d\tau)$ , where  $(d\tau, d\sigma)$  is the tangent to the curve  $\Gamma$  that we're integrating over and  $(d\sigma, -d\tau)$  gives the outward normal. So  $p_{\mu}(\Gamma) = \int_{\Gamma} (\mathcal{P}^{\tau}_{\mu} d\sigma - \mathcal{P}^{\sigma}_{\mu} d\tau)$ . The difference between some  $\Gamma$  and  $\Gamma'$  with the same endpoints (i.e.  $\partial(\Gamma - \Gamma') = 0$ ) is  $\oint_{\Gamma - \Gamma' = \partial R} (\mathcal{P}^{\tau}_{\mu} d\sigma - \mathcal{P}^{\sigma}_{\mu} d\tau) = \int_{R} d\tau d\sigma (\partial_{\tau} \mathcal{P}^{\tau}_{\mu} + \partial_{\sigma} \mathcal{P}^{\sigma}_{\mu}) = 0.$ 

• Using the  $\mathcal{P}^{\alpha\mu}$  that we found last week in static gauge we get for the conserved charges

$$p^{0} = \frac{E}{c} = \int \frac{T_{0}ds}{\sqrt{1 - v_{\perp}^{2}/c^{2}}}, \qquad \vec{p} = \int \frac{T_{0}ds}{c^{2}} \frac{v_{\perp}}{\sqrt{1 - v_{\perp}^{2}/c^{2}}}.$$

• Lorentz symmetry comes from the worldsheet symmetry  $\delta X^{\mu} = \epsilon^{\mu\nu} X_{\nu}$ , which is a symmetry if  $\epsilon^{\mu\nu} = \epsilon^{[\mu\nu]}$ , e.g.  $\delta(\eta_{\mu\nu}X^{\mu}X^{\nu}) = 0$ . The terms in the string Lagrangian  $\mathcal{L}_{NG} = -\frac{T_0}{c} \sqrt{(\dot{X} \cdot X')^2 - (\dot{X})^2 (X')^2}$  involve  $\eta_{\mu\nu} \partial_{\alpha} X^{\mu} \partial_{\beta} X^{\nu}$ , which again is invariant under  $\delta X^{\mu} = \epsilon^{\mu\nu} X_{\nu}$ . The associated conserved currents are  $\mathcal{M}^{\alpha}_{\mu\nu} = X_{\mu}\mathcal{P}^{\alpha}_{\nu} - (\mu \leftrightarrow \nu)$ . The corresponding charges  $M_{\mu\nu} = \int (\mathcal{M}^{\tau}_{\mu\nu} d\sigma - \mathcal{M}^{\sigma}_{\mu\nu} d\tau)$  are the angular momenta. Note that  $M^{0i} = ctp^i - \int d\sigma X^i \mathcal{P}^{\tau 0}$ , which can be interpreted as  $X^i_{cm}(t) = \frac{-cM^{0i}}{E} + t\frac{c^2p^i}{E}$ .