

1/8/08 Lecture outline

Recall from last quarter that physical observables (scattering probabilities, cross-sections, lifetimes, etc) can be computed from S-matrix amplitudes: observable $\sim |\langle f|S|i\rangle|^2 \cdot (\text{phase space factors})$. And the S-matrix amplitudes can be computed via Dyson's formula for the propagator, and Wick's theorem:

$$U(t_2, t_1) = T \exp(-i \int_{t_1}^{t_2} H_{int}(t') dt'),$$

$$T(\phi_1 \dots \phi_n) =: \phi_1 \dots \phi_n : + : \text{all contractions} :$$

This is nicely expressed in terms of Feynman diagrams.

- Our first topic is the Feynman path integral. Gives another way to quantize particles, and fields. For particles, consider time evolution operator

$$U(x_a, x_b; T) = \langle x_b | e^{-iHT/\hbar} | x_a \rangle.$$

Satisfies SE

$$i\hbar \partial_T U = HU.$$

Feynman:

$$U(x_a, x_b; T) = \int [dx(t)] e^{iS[x(t)]/\hbar}.$$

- Motivation. • Computation. Integral can be broken into time slices, as way to define it. E.g. free particle

$$\left(\frac{-im}{2\pi\hbar\epsilon}\right)^{N/2} \int \prod_{i=1}^{N-1} dx_i \exp\left[\frac{im}{2\hbar\epsilon} \sum_{i=1}^N (x_i - x_{i-1})^2\right]$$

Where we take $\epsilon \rightarrow 0$ and $N \rightarrow \infty$, with $N\epsilon = T$ held fixed.